



# Observations of the 18cm OH lines of 23 comets

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Presented by Aysha Rahman





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# Overview

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# Background

003-1040559 1250 003-77156.8 1760 0009-14563.7 73273



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# Motivation

- Primitive solar system bodies
- Early solar system/solar system formation
- Earth's evolution and water origin

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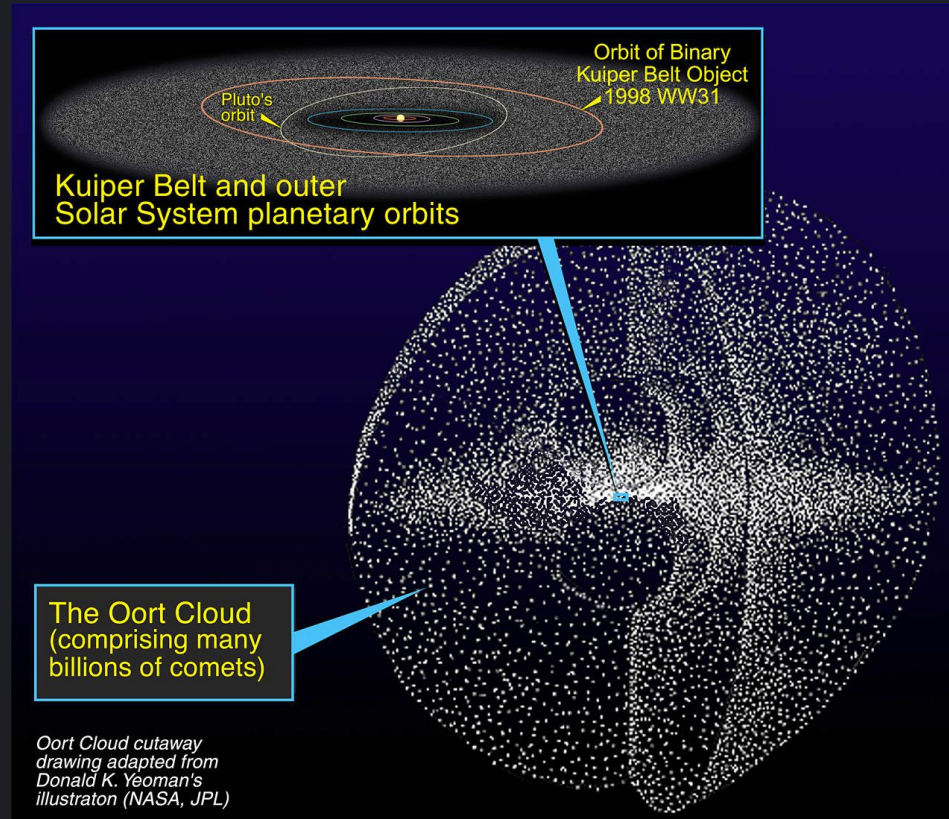
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# Orbits



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Figure 1: Oort Cloud and Kuiper Belt diagram  
Credit: NASA/ESA and A. Feild (Space Telescope Science Institute)





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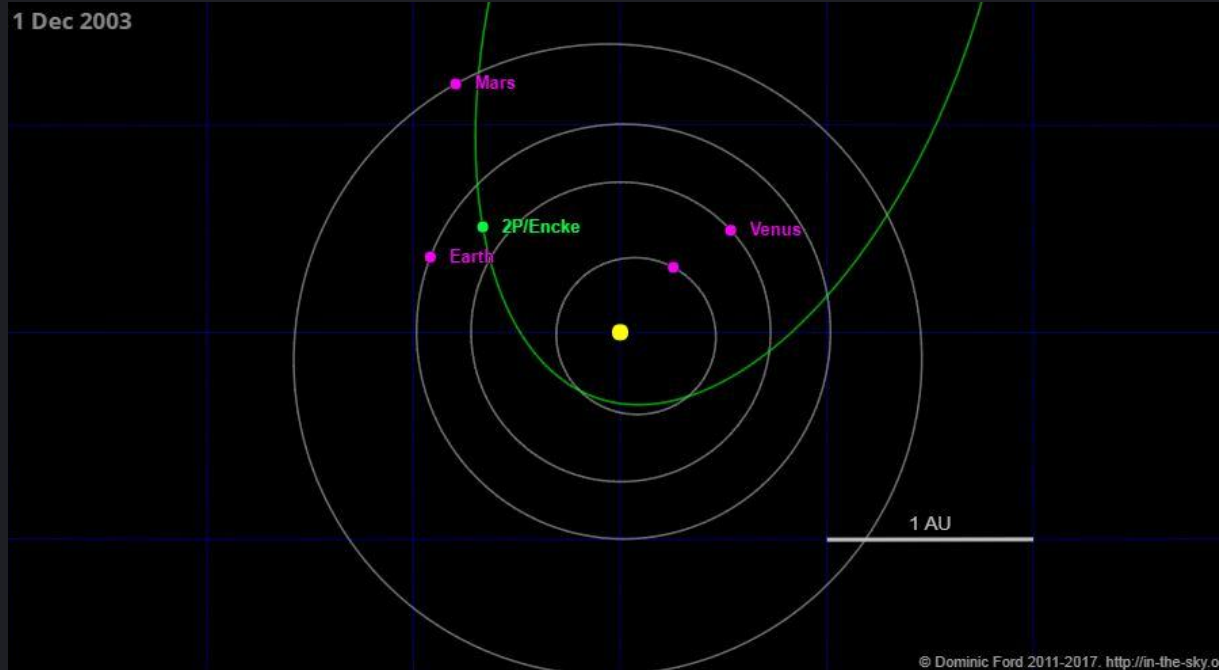
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# Sample Orbit



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Figure 2: Orbit of comet 2P/Encke  
Credit: Dominic Ford, [in-the-sky.org](http://in-the-sky.org)





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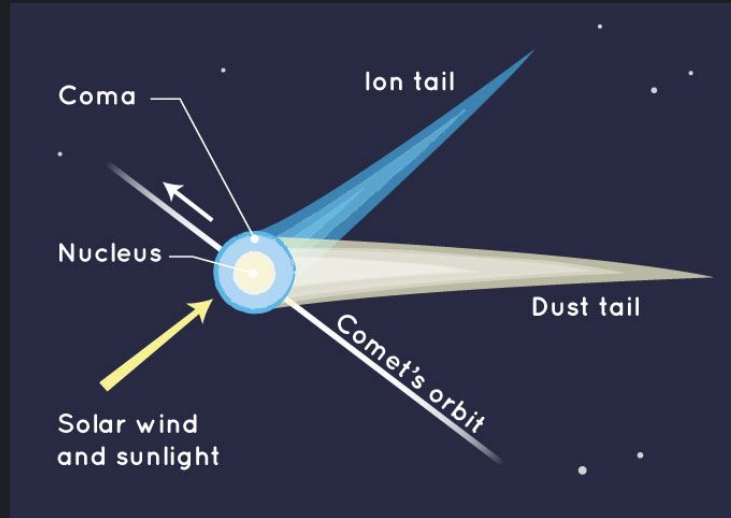
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# Comet Anatomy



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Figure 3: Diagram of comet anatomy  
Credit: NASA/JPL-Caltech



# Photodissociation

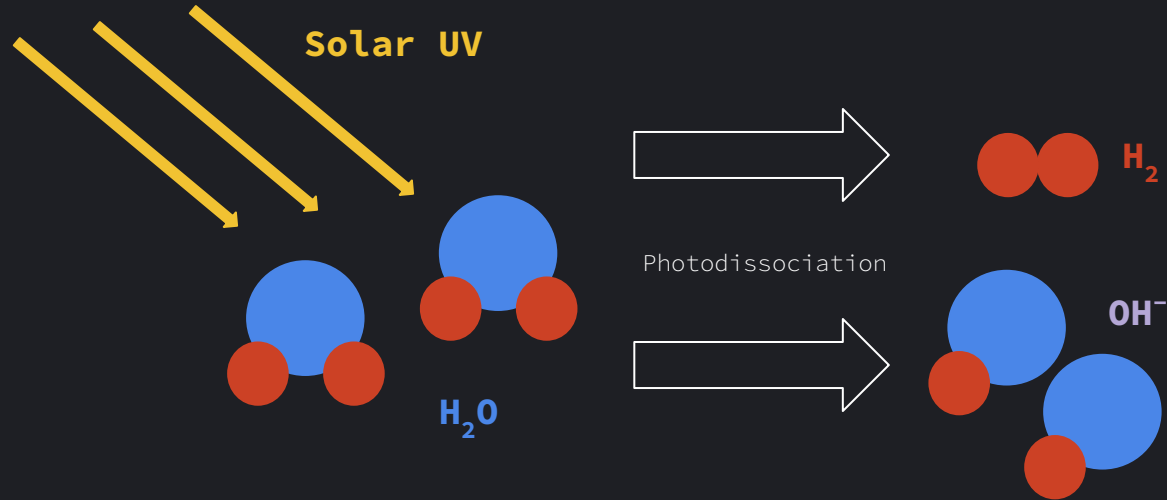


Figure 4: Photodissociation of water molecules into hydrogen and hydroxide ions





# Methods



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# Aim

- To map and model gas production rates and outflow velocities of OH<sup>-</sup> molecules in comets nearing perihelion

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# Observations

- Data spanning 16 years (2001-2017)
- Over 30 comets; we look at 23
- Arecibo Observatory and Green Bank Telescope, ground-based radio observations, 18 cm  $\Lambda$ -doublet

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# Monte Carlo Simulation

- Produce synthetic coma, divide into sections, look at doppler shift to map
- Generate  **$10^6$  H<sub>2</sub>O molecules**, sublimated
- Randomly assign parent production time, parent destruction time, daughter destruction time, and azimuthal ( $\phi$ ) and polar ( $\theta$ ) angles of motion for each molecule
- Grid of different sublimation/initial velocity  $\mathbf{v}_0$  and quenching radii  $r_q$

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# Calculations

- Parent lifetime  $\Delta t_1$ , sine and cosine projections

- Final parent position:

$$x_{12} = v_0 \Delta t_1 \sin(\Theta_1) \cos(\Phi_1)$$

$$y_{12} = v_0 \Delta t_1 \sin(\Theta_1) \sin(\Phi_1)$$

$$z_{12} = v_0 \Delta t_1 \cos(\Theta_1)$$

- Velocity due to photodissociation  $v_p = 1.05 \text{ km/s}$
- Final daughter x-velocity, sublimation + photodissociation

$$v_3 = v_0 \sin(\Theta_1) \cos(\Phi_1) + v_p \sin(\Theta_2) \cos(\Phi_2)$$



# Results

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# Spectral Mapping

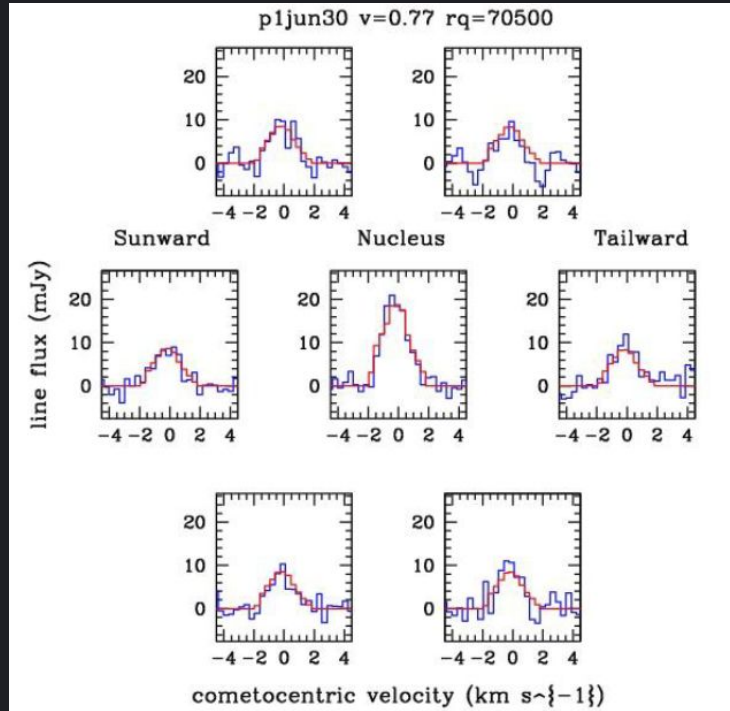
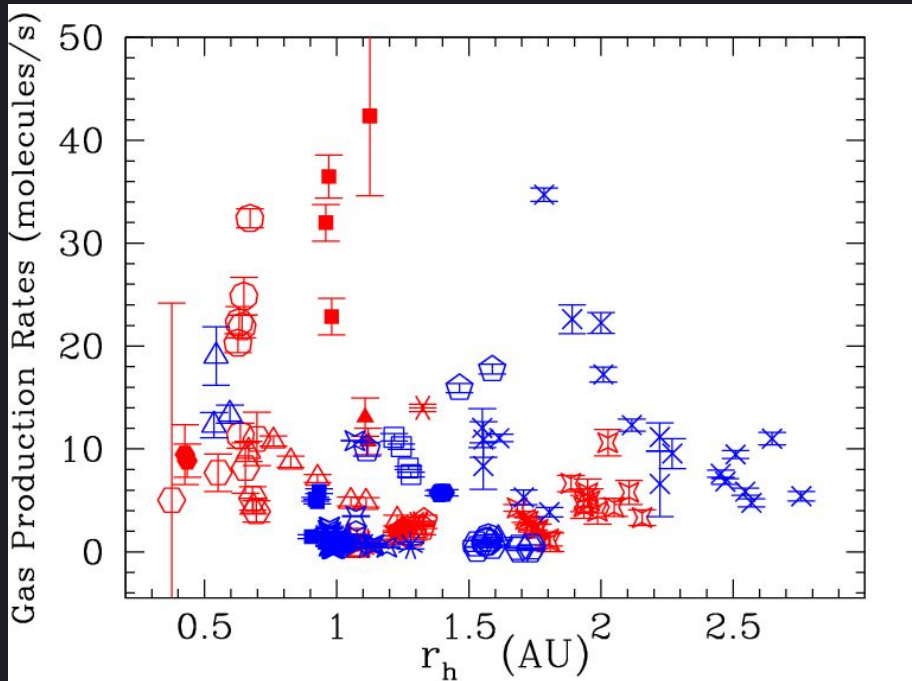


Figure: Spectrum for C/2009 P1 Garradd at 2.757 AU from the Sun on June 30, 2011

# Gas Production vs Heliocentric Distance

- ✕ C/2012 X1 LINEAR
- ◊ C/2011 L4 PANSTARRS
- C/2006 M4 SWAN
- C/2002 T7 LINEAR
- 8P Tuttle
- ◻ 103P Hartley 2
- ▲ C/2001 A2 LINEAR
- ◐ C/1999 T1 LINEAR
- C/2007 F1 LONEOS
- △ C/2007 E2 Lovejoy
- ◑ C/2013 R1 McNaught
- ✕ C/2014 Q2 Lovejoy
- ✕ C/2009 P1 Garradd
- ▲ C/2003 T4 LINEAR
- C/2007 N3 Lulin
- ◊ 9P Tempel 1
- ◻ 153P Ikeya-Zhang
- C/2007 W1 Boattini
- ✕ 73P S-W 3
- \* C/2012 S1 ISON
- ◐ C/2003 K4 LINEAR
- ◻ C/2004 Q2 Machholz
- C/2017 01 ASASSN



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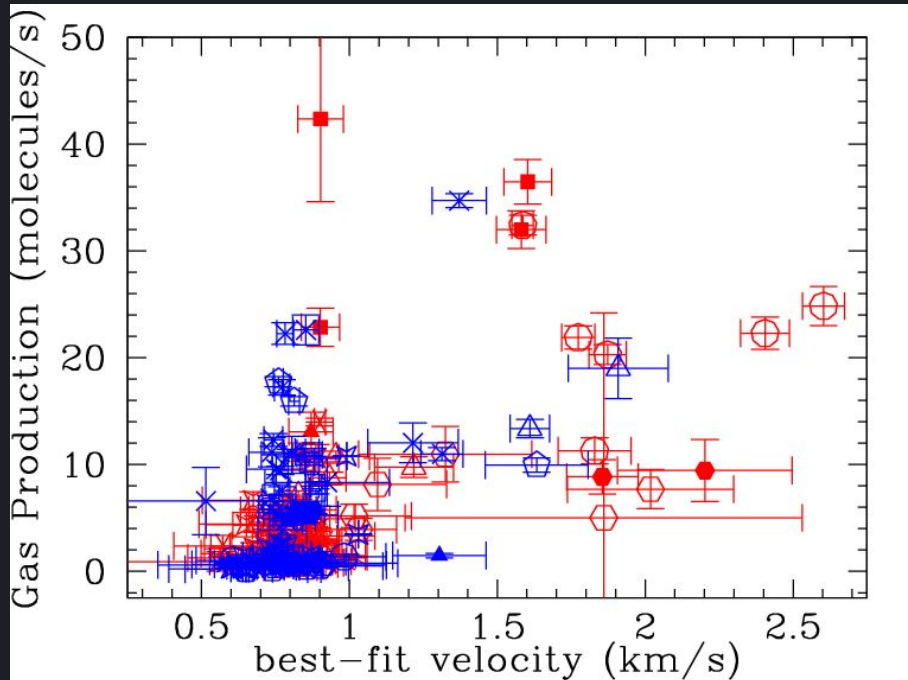
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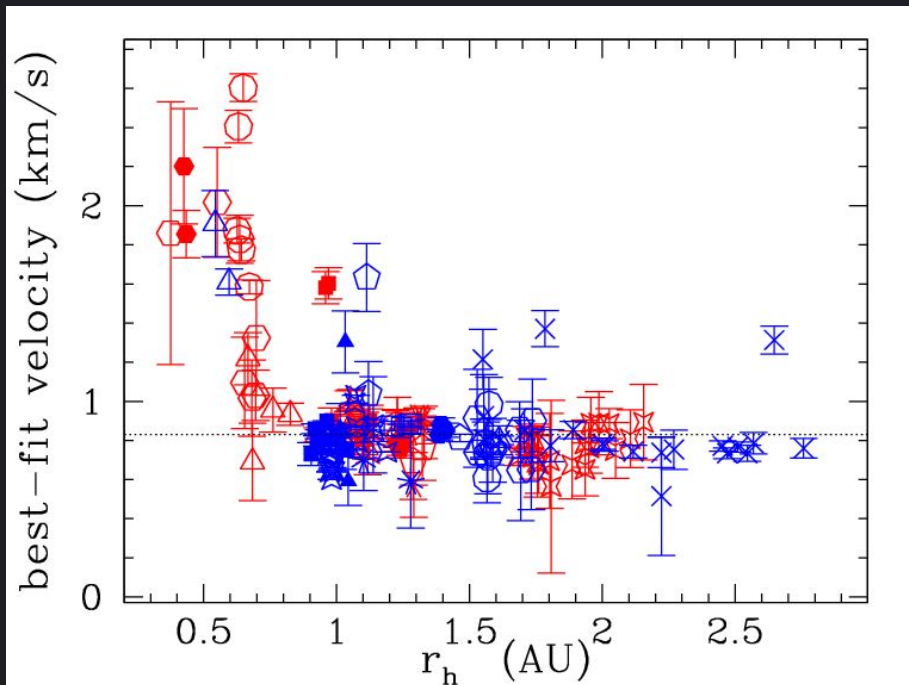
# Gas Production vs Gas Velocity

✕	C/2012 X1 LINEAR
○	C/2011 L4 PANSTARRS
■	C/2006 M4 SWAN
○	C/2002 T7 LINEAR
□	8P Tuttle
□	103P Hartley 2
▲	C/2001 A2 LINEAR
○	C/1999 T1 LINEAR
●	C/2007 F1 LONEOS
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# OH Velocity vs Heliocentric Distance

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*	C/2012 S1 ISON
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○	C/2017 O1 ASASSN





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# Conclusions and Future Work

- Correlations between gas outflow velocity and heliocentric distance
- Fewer correlations between gas production and distance or gas production and velocity
- Possible reasons for poor fit for some comets: high velocities, spectral inversion factor
- Examine different OH initial velocities

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# Sources

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- A. J. Lovell, E. S. Howell, and F. P. Schloerb (unpublished). Observations and Modeling of 18 cm OH Lines in Twenty-Seven Comets. 2016.
- A. J. Lovell, E. S. Howell, F. P. Schloerb, B. M. Lewis, and A. A. Hine. Arecibo observations of the 18cm OH lines of six comets. *Asteroids, Comets, and Meteors: ACM 2002* 500:681-684. 2002.
- A. J. Lovell, A. Bonsall, H. Sizemore, and F. Ghigo. Ground-based Observations of Small Solar System Bodies: Probing Our Local Debris Disk. *Bulletin of the AAS*, 51.3. 2019.
- A. L. Cochran, A. Levasseur-Regourd, M. Cordiner, E. Hadamcik, J. Lasue, A. Gicquel, D. G. Schliecher, S. B. Charnley, M. J. Mumma, L. Paganini, D. Bockelee-Morvan, N. Bover, and Y. Kuan. The composition of comets. *Space Science Reviews*, 197.1:9-46, July 2015.
- B. Presler-Marshall, Bachelor's thesis, Agnes Scott College, 2019.
- B. Presler-Marshall, Master's thesis, University of Central Florida, 2021.
- L. E. Tacconi-Garman, F. P. Schloerb, and H. J. Claussen. High spectral resolution observations and kinematic modeling of the 1667 MHz hyperfine transition of OH in Comets Halley (1982i), Giacobini-Zinner (1984e), Hartley-Good (1985l), Thiele (1985m), and Wilson (1986l). *The Astrophysical Journal*, 364:672-686. 1990.



# Acknowledgments



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